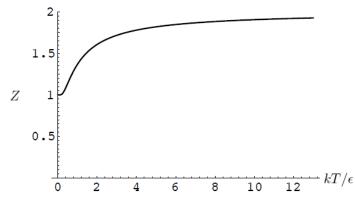
## **HW-Solution-set7**

## 6.3, 6.5, 6.10, 6.12, 6.13, 6.16, 6.25, 6.29, 6.31, 6.33, 6.37, 6.44, 6.48, 6.49, 6.52.

Problem 6.3. For this two-state system,

$$Z = e^0 + e^{-\epsilon/kT} = 1 + e^{-\epsilon/kT},$$

where  $\epsilon=2$  eV is the energy of the excited state. Thus Z varies between 1 (in the limit  $T\to 0$ ) and 2 (in the limit  $T\to \infty$ ). To plot the partition function vs. temperature it's convenient to define the dimensionless variable  $t=kT/\epsilon$ , so  $Z=1+e^{-1/t}$ . The value of t at T=300 K is about 1/80=.013, while the value of t at T=300,000 K is about 13. (Recall that at 300 K, kT=.026 eV  $\approx 1/40$  eV.) Here, then, is a plot of Z vs. t for values of t up to 13:



Plugging in the particular temperatures given yields the following values:

$$T = 300 \text{ K}: \qquad Z = 1 + e^{-1/.013} = 1 + 2.2 \times 10^{-34}$$

$$T = 3000 \text{ K}: \qquad Z = 1 + e^{-1/.13} = 1.00043$$

$$T = 30,000 \; {\rm K}: \qquad Z = 1 + e^{-1/1.3} = 1.46$$

$$T = 300,000 \text{ K}$$
:  $Z = 1 + e^{-1/13} = 1.93$ 

Notice that the approach to 1 at low temperature is much more dramatic than the approach to 2 at high temperature.

Problem 6.5. (A three-state toy model.)

(a) At 300 K, kT = 0.026 eV, as computed on page 13. Therefore the partition function for this system is

$$Z = e^{-(-0.05/0.026)} + e^{0} + e^{-(0.05/0.026)} = 6.84 + 1 + 0.15 = 7.99.$$

(b) Numbering the states 1, 2, and 3 in the order listed, the probabilities are

$$\mathcal{P}_1 = \frac{6.84}{7.99} = 0.86;$$
  $\mathcal{P}_2 = \frac{1}{7.99} = 0.13;$   $\mathcal{P}_3 = \frac{0.15}{7.99} = 0.02.$ 

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(c) Measuring the energies now relative to the ground state, we have

$$Z = e^{0} + e^{-(0.05/0.026)} + e^{-(0.10/0.026)} = 1 + 0.15 + 0.02 = 1.17.$$

And the probabilities are

$$\mathcal{P}_1 = \frac{1}{1.17} = 0.86;$$
  $\mathcal{P}_2 = \frac{0.15}{1.17} = 0.13;$   $\mathcal{P}_3 = \frac{0.02}{1.17} = 0.02.$ 

So even though the partition function changes, the probabilities are unchanged, as they must be because nature can't possibly care what we use as our zero-point for measuring energy.

## Problem 6.10. (Vibrational excitations of H<sub>2</sub>O.)

(a) The partition function for this vibrating atom is

$$Z = e^{-hf/2kT} + e^{-3hf/2kT} + e^{-5hf/2kT} + \cdots$$

At 300 K,

$$\frac{hf}{kT} = \frac{(4.136 \times 10^{-15} \text{ eV} \cdot \text{s})(4.8 \times 10^{13} \text{ s}^{-1})}{(8.617 \times 10^{-5} \text{ eV/K})(300 \text{ K})} = 7.68,$$

so the partition function is approximately

$$Z = e^{-3.84} + e^{-11.52} + e^{-19.20} + \cdots$$
  
= 0.0215 + (9.9 × 10<sup>-6</sup>) + (4.6 × 10<sup>-9</sup>) + \cdots  
= 0.0215.

The probabilities of the lowest two excited states are therefore

$$\mathcal{P}_1 = \frac{9.9 \times 10^{-6}}{0.0215} = 0.00046, \qquad \mathcal{P}_2 = \frac{4.6 \times 10^{-9}}{0.0215} = 2.1 \times 10^{-7}.$$

The probability of the ground state is very nearly 1; more precisely,

$$\mathcal{P}_0 = 1 - 0.00046 = 0.99954.$$

(b) At 700 K, the ratio hf/kT is smaller by a factor of 3/7, so hf/kT=3.25 and the partition function is

$$Z = e^{-1.624} + e^{-4.873} + e^{-8.121} + e^{-11.37} + \cdots$$
$$= 0.1971 + 0.0077 + 0.0003 + 0.00001 + \cdots$$
$$= 0.2051.$$

The probabilities are therefore

$$\mathcal{P}_0 = \frac{0.1971}{0.2051} = 0.961, \qquad \mathcal{P}_1 = \frac{0.0077}{0.2051} = 0.038, \qquad \mathcal{P}_2 = \frac{0.0003}{0.2051} = 0.001.$$

**Problem 6.12.** If the molecules are in equilibrium with a reservoir of temperature T, then the probability of a molecule being in any one of the excited states, relative to the ground state, should be

$$\frac{e^{-E_1/kT}}{e^{-E_0/kT}} = e^{-(E_1 - E_0)/kT}.$$

We are given that this relative probability is approximately 1/10. Therefore,

$$-\frac{E_1 - E_0}{kT} = \ln \frac{1}{10} = -2.303,$$

or

$$T = \frac{4.7 \times 10^{-4} \text{ eV}}{(2.303)(8.62 \times 10^{-5} \text{ eV/K})} = 2.4 \text{ K}.$$

The uncertainty in the data, however, is somewhat large. We now know that the temperature is closer to 2.7 K, and that the "reservoir" is the cosmic background radiation, a gas of photons that fills the entire observable universe (see Section 7.4).

Problem 6.13. The ratio of probabilities under these conditions should be

$$\frac{\mathcal{P}(n)}{\mathcal{P}(p)} = \frac{e^{-m_n c^2/kT}}{e^{-m_p c^2/kT}} = e^{-(\Delta m)c^2/kT} = \exp\left(-\frac{(2.3 \times 10^{-30} \text{ kg})(3 \times 10^8 \text{ m/s})^2}{(1.38 \times 10^{-23} \text{ J/K})(10^{11} \text{ K})}\right) = 0.86.$$

In other words, there should be 86 neutrons for every 100 protons. That makes 186 particles total, so the fraction of protons should be 100/186 = 0.54, and the fraction of neutrons should be 86/186 = 0.46.

**Problem 6.16.** Starting from the definition of Z,

$$\frac{\partial Z}{\partial \beta} = \frac{\partial}{\partial \beta} \sum_{s} e^{-\beta E(s)} = \sum_{s} \frac{\partial}{\partial \beta} e^{-\beta E(s)} = \sum_{s} (-E(s)) e^{-\beta E(s)}.$$

Now just multiply by -1/Z, cancel the minus signs, and move the Z inside the sum:

$$-\frac{1}{Z}\frac{\partial Z}{\partial \beta} = -\frac{1}{Z}\sum_{s}(-E(s))e^{-\beta E(s)} = \sum_{s}E(s)\frac{e^{-\beta E(s)}}{Z} = \sum_{s}E(s)\mathcal{P}(s) = \overline{E}.$$

**Problem 6.25.** Because of the symmetry of the CO<sub>2</sub> molecule, a 180° rotation has no effect on its state, and therefore we should divide its rotational partition function by 2 just as for a diatomic molecule with identical atoms:

$$Z_{\text{rot}} = \frac{kT}{2\epsilon} = \frac{0.026 \text{ eV}}{2(0.000049 \text{ eV})} = 265.$$

**Problem 6.29.** From the graph plotted in the previous problem, we see that the rotational heat capacity falls off steeply when  $kT/\epsilon$  is between about 0.3 and 0.6; the heat capacity is at about half its asymptotic value when  $kT/\epsilon \approx 0.4$ . For HD, that translates to a temperature of

$$T = \frac{0.4 \,\epsilon}{k} = \frac{(0.4)(0.0057 \text{ eV})}{8.62 \times 10^{-5} \text{ eV/K}} = 26 \text{ K}.$$

**Problem 6.31.** If state q has energy c|q|, then the Boltzmann factor for this state is  $e^{-\beta c|q|}$  and the partition function is the sum of all these Boltzmann factors:

$$Z = \sum_{q} e^{-\beta c|q|} = \frac{1}{\Delta q} \sum_{q} e^{-\beta c|q|} \Delta q.$$

In the limit  $\Delta q \to 0$  the sum becomes an integral:

$$Z = \frac{1}{\Delta q} \int_{-\infty}^{\infty} e^{-\beta c|q|} dq = 2 \cdot \frac{1}{\Delta q} \int_{0}^{\infty} e^{-\beta cq} dq = \frac{2}{\Delta q} \left( -\frac{1}{\beta c} \right) e^{-\beta cq} \Big|_{0}^{\infty} = \frac{2}{\beta c \Delta q} \equiv C\beta^{-1}.$$

(Since the integral is symmetric under  $q \to -q$ , I've written it as twice the integral over only positive q values. This trick gets rid of the absolute value bars.) The average energy is therefore

$$\overline{E} = -\frac{1}{Z} \frac{\partial Z}{\partial \beta} = -\frac{\beta}{C} (-C\beta^{-2}) = \frac{1}{\beta} = kT.$$

Problem 6.33. The mass of an oxygen molecule is 32 u, so for oxygen at 300 K,

$$\sqrt{\frac{kT}{m}} = \sqrt{\frac{(1.38 \times 10^{-23} \text{ J/K})(300 \text{ K})}{32(1.66 \times 10^{-27} \text{ kg})}} = 279 \text{ m/s}.$$

The most likely speed,  $v_{\text{max}}$ , is just this times  $\sqrt{2}$ , or 395 m/s. To get the rms speed, we instead multiply by  $\sqrt{3}$  to get 484 m/s. And to get the average speed, we multiply by  $\sqrt{8/\pi}$  to get 446 m/s.

**Problem 6.37.** In analogy with equation 6.51, imagine first that the v values are discretely spaced, then take the continuum limit:

$$\overline{v^2} = \sum_{v} v^2 \mathcal{D}(v) \, dv \to \int_0^\infty v^2 \mathcal{D}(v) \, dv = \frac{4}{\sqrt{\pi}} \left(\frac{m}{2kT}\right)^{3/2} \int_0^\infty v^4 e^{-mv^2/2kT} \, dv$$
$$= \frac{4}{\sqrt{\pi}} \left(\frac{m}{2kT}\right)^{3/2} \left(\frac{2kT}{m}\right)^{5/2} \int_0^\infty x^4 e^{-x^2} \, dx = \frac{8kT}{\sqrt{\pi} m} \int_0^\infty x^4 e^{-x^2} \, dx.$$

The integral over x is worked out in Appendix B, Problem B.2; or you can look it up in a table or ask a computer. The answer is  $3\sqrt{\pi}/8$ , so finally we have

$$\overline{v^2} = \frac{8kT}{\sqrt{\pi} \, m} \cdot \frac{3\sqrt{\pi}}{8} = \frac{3kT}{m},$$

in agreement with equation 6.41 and the equipartition theorem.

**Problem 6.44.** For N indistinguishable, noninteracting molecules that can exchange places with each other,

$$Z = \frac{1}{N!} Z_1^N,$$

 $\mathbf{SO}$ 

$$F = -kT \ln Z = -kT \left[ N \ln Z_1 - \ln N! \right]$$
$$= -kT \left[ N \ln Z_1 - N \ln N + N \right] = -NkT \left[ \ln \frac{Z_1}{N} + 1 \right].$$

Therefore the chemical potential is

$$\mu = \left(\frac{\partial F}{\partial N}\right)_{T,V} = -kT\left[\ln\frac{Z_1}{N} + 1\right] - NkT\frac{\partial}{\partial N}(-\ln N) = -kT\ln\frac{Z_1}{N}.$$

**Problem 6.48.** (S and  $\mu$  for a diatomic gas.)

(a) For a collection of N rotating diatomic molecules, the internal contribution to the free energy is

$$F_{\rm int} = -NkT \ln(Z_e Z_{\rm rot}),$$

where (at ordinary temperatures)  $Z_{\rm rot}$  is  $kT/\epsilon$  for a molecule composed of two different atoms, or  $kT/2\epsilon$  for a molecule composed of identical atoms. Either way,  $Z_{\rm rot}$  is simply a constant times T, so

$$\frac{\partial F_{\text{int}}}{\partial T} = -Nk \ln(Z_e Z_{\text{rot}}) - NkT \frac{1}{T} = -Nk [\ln(Z_e Z_{\text{rot}}) + 1].$$

Therefore, according to equation 6.92,

$$S = Nk \left[ \ln \left( \frac{V}{Nv_Q} \right) + \frac{5}{2} \right] + Nk \left[ \ln \left( Z_e Z_{\text{rot}} \right) + 1 \right] = Nk \left[ \ln \left( \frac{V Z_e Z_{\text{rot}}}{Nv_Q} \right) + \frac{7}{2} \right].$$

The rotational partition function for oxygen at room temperature is

$$Z_{\text{rot}} = \frac{kT}{2\epsilon} = \frac{(8.617 \times 10^{-5} \text{ eV/K})(298 \text{ K})}{2(.00018 \text{ eV})} = 71,$$

while the quantum volume is

$$v_Q = \left(\frac{h}{\sqrt{2\pi mkT}}\right)^3 = \left(\frac{6.63 \times 10^{-34} \text{ J} \cdot \text{s}}{\sqrt{2\pi (32)(1.66 \times 10^{-27} \text{ kg})(1.38 \times 10^{-23} \text{ J/K})(298 \text{ K})}}\right)^3$$
$$= (1.79 \times 10^{-11} \text{ m})^3 = 5.73 \times 10^{-33} \text{ m}^3$$

and the average volume per particle (at atmospheric pressure) is

$$\frac{V}{N} = \frac{kT}{P} = \frac{(1.38 \times 10^{-23} \text{ J/K})(298 \text{ K})}{1.01 \times 10^5 \text{ N/m}^2} = 4.07 \times 10^{-26} \text{ m}^3.$$

From these numbers we can compute the logarithm

$$\ln\left(\frac{VZ_eZ_{\text{rot}}}{Nv_O}\right) = \ln\left(\frac{(4.07 \times 10^{-26} \text{ m}^3)(3)(71)}{5.73 \times 10^{-33} \text{ m}^3}\right) = 21.14.$$

Thus the entropy under these conditions is

$$S = Nk[21.14 + 3.50] = (24.6)nR = 205 \text{ J/K},$$

precisely in agreement with the measured value (to the number of significant figures used in the calculation).

(b) The chemical potential is -kT times the same logarithm:

$$\mu = -(8.62 \times 10^{-5} \text{ eV/K})(298 \text{ K})(21.1) = -0.54 \text{ eV}.$$

**Problem 6.49.** As shown in Section 6.2, the rotational energy of a diatomic molecule at room temperature is kT, corresponding to two degrees of freedom. Therefore the total thermal energy of a mole of  $N_2$  is

$$U = \frac{3}{2}NkT + NkT = \frac{5}{2}NkT = \frac{5}{2}nRT = \frac{5}{2}(1 \text{ mol})(8.31 \text{ J/mol} \cdot \text{K})(298 \text{ K}) = 6190 \text{ J}.$$

The enthalpy is just U + PV = U + nRT, so it's larger by

$$nRT = (1 \text{ mol})(8.31 \text{ J/mol} \cdot \text{K})(298 \text{ K}) = 2480 \text{ J}, \text{ that is, } H = 8670 \text{ J}.$$

To compute the remaining quantities we need the internal partition function, which in this case is purely rotational:

$$Z_{\text{int}} = Z_{\text{rot}} = \frac{kT}{2\epsilon} = \frac{(8.617 \times 10^{-5} \text{ eV/K})(298 \text{ K})}{2(.00025 \text{ eV})} = 51.$$

We also need the quantum volume,

$$v_Q = \left(\frac{h}{\sqrt{2\pi mkT}}\right)^3 = \left(\frac{6.63 \times 10^{-34} \text{ J} \cdot \text{s}}{\sqrt{2\pi (28)(1.66 \times 10^{-27} \text{ kg})(1.38 \times 10^{-23} \text{ J/K})(298 \text{ K})}}\right)^3$$
$$= (1.91 \times 10^{-11} \text{ m})^3 = 6.98 \times 10^{-33} \text{ m}^3,$$

and the average volume per particle,

$$\frac{V}{N} = \frac{kT}{P} = \frac{(1.38 \times 10^{-23} \text{ J/K})(298 \text{ K})}{1.01 \times 10^5 \text{ N/m}^2} = 4.07 \times 10^{-26} \text{ m}^3.$$

From these numbers we can compute the logarithm

$$\ln\left(\frac{VZ_{\text{int}}}{Nv_Q}\right) = \ln\left(\frac{(4.07 \times 10^{-26} \text{ m}^3)(51)}{6.98 \times 10^{-33} \text{ m}^3}\right) = 19.5.$$

The Helmholtz free energy is therefore

$$F = -nRT \left[ \ln \left( \frac{VZ_{\text{int}}}{Nv_Q} \right) + 1 \right] = -(2480 \text{ J})[19.5 + 1] = -50.8 \text{ kJ},$$

while the Gibbs free energy is

$$G = F + PV = -50.8 \text{ kJ} + 2480 \text{ J} = -48.3 \text{ kJ}.$$

The easiest way to get the entropy is from the definition F = U - TS:

$$S = \frac{U - F}{T} = \frac{(6190 \text{ J}) - (-50,800 \text{ J})}{298 \text{ K}} = 191 \text{ J/K}$$

(in agreement with the measured value tabulated on page 405). And the easiest way to get the chemical potential is from  $G = N\mu$ :

$$\mu = \frac{G}{N} = \frac{-48.3 \text{ kJ}}{6.02 \times 10^{23}} = 8.03 \times 10^{-20} \text{ J} = -.501 \text{ eV}.$$

**Problem 6.52.** As in the nonrelativistic case, the allowed wavelengths (in one dimension) are  $\lambda_n = 2L/n$ , and therefore the allowed momenta are  $p_n = h/\lambda_n = hn/2L$ . Now, however, the relation between energy and momentum is E = pc, so the allowed energies are  $E_n = hcn/2L$ . Therefore the single-particle partition function is

$$Z_{1d} = \sum_{n} e^{-E_n/kT} = \sum_{n} e^{-hcn/2LkT}.$$

When L is macroscopic the number of terms in the sum that are significant is very large, so we can convert the sum to an integral to obtain

$$Z_{1d} = \int_0^\infty e^{-hcn/2LkT} dn = -\frac{2LkT}{hc} e^{-hcn/2LkT} \bigg|_0^\infty = \frac{2LkT}{hc}.$$

As expected, the partition function is directly proportional to L and increases with increasing temperature.