## Light Front (LF) Nuclear Structure Gerald A Miller, U. of Washington

- formulate the NN interaction on the light front
- solve the Weinberg equation for the deuteron
- prescription from FS 81 review constructs LF wave function from NR wf:
- how good is this approx at recoil momenta few hundred MeV?
- can we get the LF wf from NN potentials?
- Nuclear Many Body LF wave function inc. Correlations -Miller & Machleidt PRC60, 035202
- Relevance of LF wave functions for quasi-elastic scattering

stimulated by Ellie Long Proposal on Tensor asymmetries work with John Terry REU J R Cooke nucl-th/0112029, Cooke & Miller PRC66, 034002 Miller, Prog. Nuc. Part. Phys. 45, 83 Tiburzi & Miller PRC81,035201 Light front quantization, Infinite momentum frame

"Time",  $x^+ = x^0 + x^3$ , "Evolve",  $p^- = p^0 - p^3$ "Space",  $x^- = x^0 - x^3$ , "Momentum",  $p^+$ (Bjorken) Transverse position, momentum  $\mathbf{b}, \mathbf{p}$   $p^- = \frac{\mathbf{p}^2 + m^2}{n^+}$ 



### Light front quantization, Infinite momentum frame

 $P^-$  is LF Hamiltonian, get from Lagrangian. LF Schroedinger eq.  $P^-|\Psi_D\rangle=M_D|\Psi_D\rangle$  Rest frame One boson exchange



Weinberg equation= Lippmann-Schwinger eq with extra factor (~1) in Green's function

Miller & Machleidt PRC 60,035202  $\pi, \eta, \rho, \omega, a_0, \sigma$  exchange with extra factor in G **NN** scattering Nuclear Matter Saturation 60  $^{1}S_{0}$ <sup>3</sup>P<sub>1</sub> Ω 40 δ (deg) δ (deg) -10 0 20 -20 0 -30200 100 300 100 200 300 0 0  $T_{lab}$  (MeV)  $T_{lab}$  (MeV) (d) (a) 20 <sup>3</sup>P<sub>0</sub>  ${}^{3}S_{1}$ E/A (MeV) 120 -10 10 δ (deg) δ (deg) 80 40 0 -10200 300 200 100 0 100 300 0  $T_{lab}$  (MeV)  $T_{lab}$  (MeV) (e) (b) -2010 10 <sup>3</sup>D<sub>1</sub>  $^{1}P_{1}$ 0 0  $\delta (deg)$  $\delta$  (deg) -10 -101.5 2 -20 -20 1 -30-30  $k_{\rm F} \,({\rm fm}^{-1})$ 200 100 100 200 300 0 300 0  $T_{lab}$  (MeV)  $T_{lab}$  (MeV) (f) (c)

 $B_{\rm D} = 2.245 \,\mathrm{MeV}$ 

## The real problem- Bethe Salpeter Eq. (BSE)

K is sum of irred.

diagrams

G depends on 4-momenta-product of two Feynman propagators

 $T = \mathbf{X} + \mathbf{X}$ 

Reduce to 3 dimensions:

ET: integrate over  $k^0$ . Ignore  $k^0$  except in G. Sets relative time to 0. LF: Integrate over  $k^-$ . Ignore  $k^-$  except in G. Sets relative  $\tau = 0$ 3 dimensional version of G is  $g_{ET}$  (Blankenbecler Sugar) or  $g_{LF}$  (Weinberg) **Puts BOTH particles on mass shell** 

No relation between wave functions in principle Spectator on-shell- Gross equation- one particle off mass shell  $k^2 - m^2 \neq 0$ 

#### PHYSICAL REVIEW C 81, 035201 (2010)

#### **Relation between equal-time and light-front wave functions**

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The relation between equal-time and light-front wave functions is studied using models for which the fourdimensional solution of the Bethe-Salpeter wave function can be obtained. The popular prescription of defining the longitudinal momentum fraction using the instant-form free kinetic energy and third component of momentum is found to be incorrect except in the nonrelativistic limit. One may obtain light-front wave functions from rest-frame, instant-form wave functions by boosting the latter wave functions to the infinite momentum frame.

- How bad is the problem?
- Is D non-relativistic?
- Is <sup>3</sup>He non-relativistic?
- Answer by using solutions of Bethe-S eqn.



### FIG. 2. Bethe-Salpeter equation for a point interaction. The state is bound by the infinite chain of bubbles.

### Model



## <sup>3</sup>He Compare true LFD from BSE with ansatz from ET



# Are light front or equal time wave functions relevant?

- Light front wave functions on-mass shell  $k_i^2 = m^2$
- Same for equal time wave functions Mismatch!
- Answer this question through explicit example point interaction model used above
- Space-like Form factors -use of light front wave functions gives exact results- GAM Phys.Rev. C80 (2009) 045210
- What about quasi-elastic scattering?

### quasi-elastic scattering



2. On- shell  $k^2 = m^2$  in 4-momentum conservation delta function  $\frac{Q^2}{\nu^2} << 1$ 

3. As in 2, Bjorken limit







# Low Q, x=I (Long's )exp't



On-mass shell approximation V=0 is OK can us standard wave functions

# Low Q, x=1.5



Significant error by assuming on shellcan't use standard wave functions spectator on-shell is ok

# Low Q, x=0.5



Significant error by assuming on shell can't use standard wave functions spectator on-shell is ok

# Non-relativistic approximation



# Summary

- formulate the NN interaction on the light front get from BSE
- solve Weinberg equation for the deuteron -done
- prescription from FS 81 review constructs LF wave function from NR wf: study with exact solutions of BSE
- how good is this approx at recoil momenta few hundred MeV?-seems ok up to about 250-300 MeV more study needed
- can we get the LF wf from NN potentials?-seems ok
- Relevance of LF wave functions- OK for space-like form factors
- Quasi-elastic at Q^2< I5 GeV^2, LF wave function No Good, Need spectator wf if x is not very near I
- Non-relativistic approximation is not good except in narrow range of x, depends on k\_perp

# Spares follow

The APS Council and the DNP have endorsed the establishment of the

#### Herman Feshbach Prize in Nuclear Physics

Purpose: To recognize and encourage outstanding research in theoretical nuclear physics. The prize will consist of \$10,000 and a certificate citing the contributions made by the recipient. The prize will be presented biannually or annually.

Herman Feshbach was a dominant force in Nuclear Physics for many years. The establishment of this prize depends entirely on the contributions of institutions, corporations and individuals associated with Nuclear Physics. So far, significant contributions have been made by MIT, the DNP, ORNL/U.Tenn, JSA/SURA, BSA, Elsevier Publishing, TUNL, TRIUMF, MSU, and a number of individuals. More than \$150,000 has been raised, primarily through institutional contributions. **It is very important that physicists make contributions to carry the endowment over the \$200,000 mark, so that the Prize will be eligible to be awarded annually.** Please help us reach that goal by making a contribution. Go online at <u>http://www.aps.org/</u> Look for the support banner and click APS member (membership number needed) and look down the list of causes.

If you have any questions, please contact G. A. (Jerry) Miller UW, <miller@uw.edu>.

Magnetic field (G)

II



(Hg) 1E-13

### If annual- number of experimentalists winning Bonner prize goes up by >50%





Extra factor is close to unity for D wave function

# Jason Cooke nucl-th/0112029, Cooke & Miller PRC66, 034002

Solves LF Schroedinger eq (LFSSE)

 $\left|P_{0}^{-}+V(P^{-})\right|\left|\Psi_{D}\right\rangle=P^{-}\left|\Psi_{D}\right\rangle$  $P^- = 2m - B$  rest frame  $V(P^{-}) = \begin{bmatrix} \mathbf{T} & \text{Planifest rotational invar. Droker} \\ \mathbf{T} & = g^2 \frac{1}{P^{-} - k_3^{-} - k_2^{-} - k_{\pi}^{-}} \\ \text{Different meson propagator than Machleidt Miller} \end{bmatrix}$ Manifest rotational invar. broken **2** Solve LSSE using transformation from  $\alpha$  to  $k_z$ :  $\alpha = \frac{k^+}{P^+} = \frac{k^+}{2M - B} = \frac{1}{2} \frac{\sqrt{\vec{k}^2 + m^2 + k_z}}{\sqrt{\vec{k}^2 + m^2}}$ Solve w. rot. inv. in  $\perp$  plane (polar coords) Computed B depends on magnetic quantum number!

## Cooke nucl-th/0112029, Cooke & Miller PRC66, 034002 Dynamics

- Chiral Lagrangian with  $\pi, \eta, \rho, \omega, \delta, \sigma$
- Two meson exchange!
- Explicit P<sup>-</sup> dependence



FIG. 1. The first several terms of the full kernel for the Bethe-Salpeter equation of the nuclear model with chiral symmetry.

## Cooke nucl-th/0112029, Cooke & Miller PRC66, 034002 Two Meson Dynamics





FIG. 9. The values of the binding energy for the m=0 and m = 1 states for different nucleon-nucleon light-front potentials. The  $\sigma$ 

# Restoring RI in form factors

- Rotational invariance gives angular condition FS
- Angular condition is upheld better when Deut is computed using only one meson exchange OME potentials than two meson exchange TME
- However, form factors do not depend much on choice of bad currents



# Miller & Machleidt PRC 60,035202 Nuclear Matter Saturation



### Solid -our light front Dashed- ET formalism